A Stellar Age Dependence of the Planet Radius Gap

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The Planet Radius Gap



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Two Competing Theories:

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1. Photoevaporation

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2. Core-powered mass-loss

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<u>Models</u>

Observations



Owen & Wu (2017)

Fulton et al. (2017)

2. Core-powered mass-loss

<u>Models</u>

Observations



Gupta & Schlichting (2019)

Fulton & Petigura (2018)

How do we differentiate between the two theories if both describe the observations well so far?









Age Validation: Clusters



Age Validation: Asteroseismology

Now that we have these shiny new stellar properties, what can we learn about planets?

How does the planet radius valley depend on stellar age?

Summary/Conclusions

- The Gaia-Kepler Stellar Properties Catalog (Berger et al. 2020a) has a median fractional radius precision of ~4%, compared to 25% for the previous Kepler Catalog.
- We validated our isochrone ages (median catalog fractional uncertainty ~56%) with cluster and asterseismic comparisons, which agree within reported uncertainties.
- We find first evidence for the planet radius gap's dependence on stellar age on timescales of a Gyr.