

Test of Gravity with NS Mergers

Kent Yagi

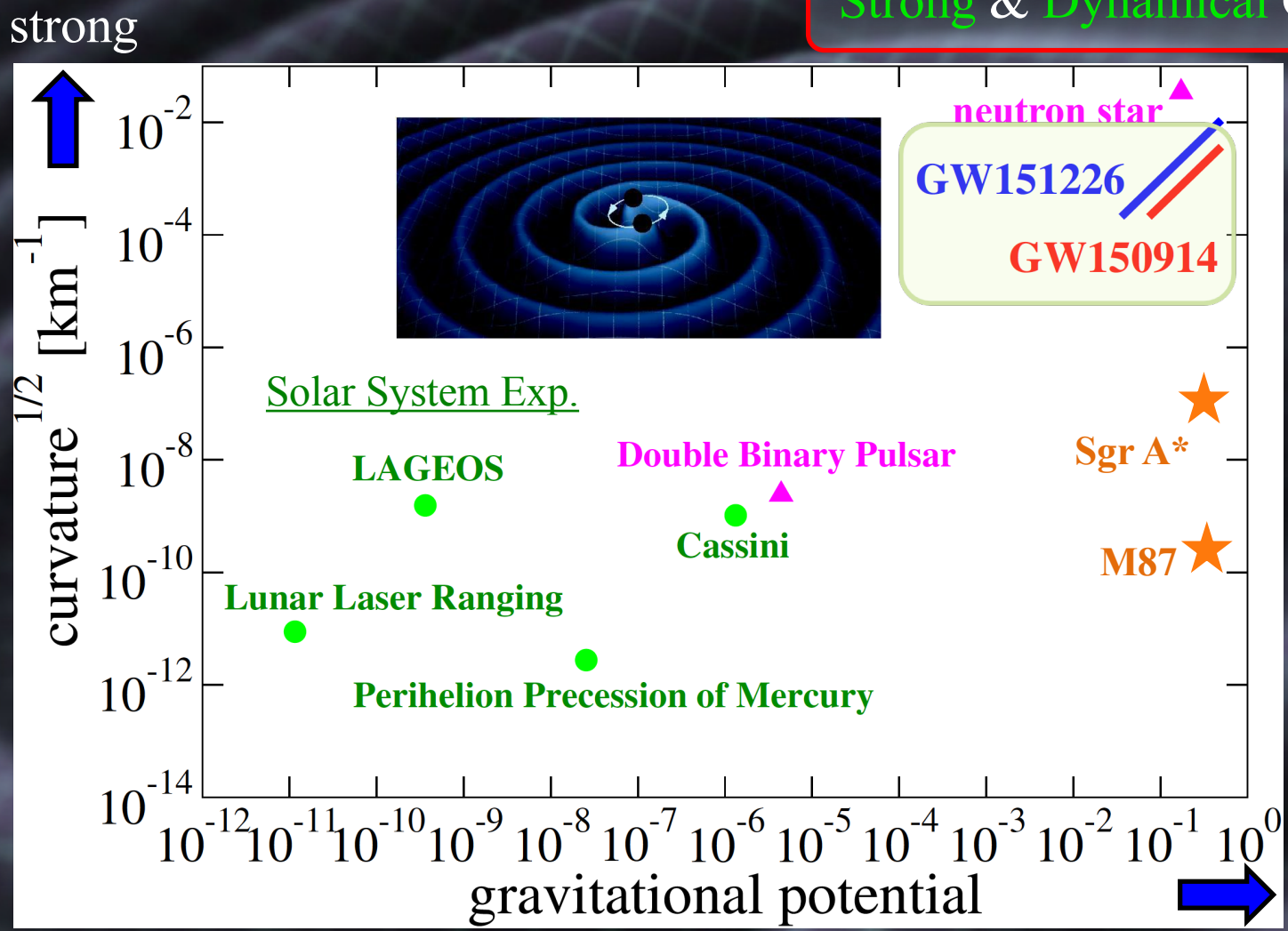
University of Virginia

KITP, Santa Barbara

June 26th 2019

Sources for Tests of Gravity

Strong & Dynamical Gravity

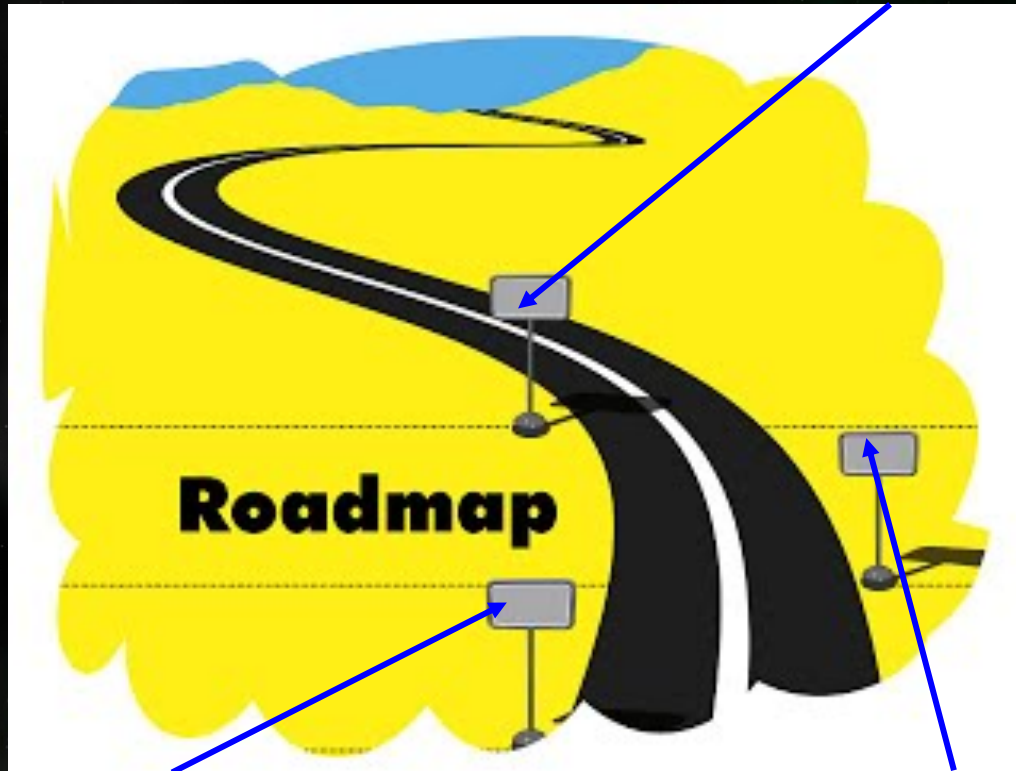


[Yunes, KY & Pretorius PRD (2016)]

large

Roadmap

Universal Relations



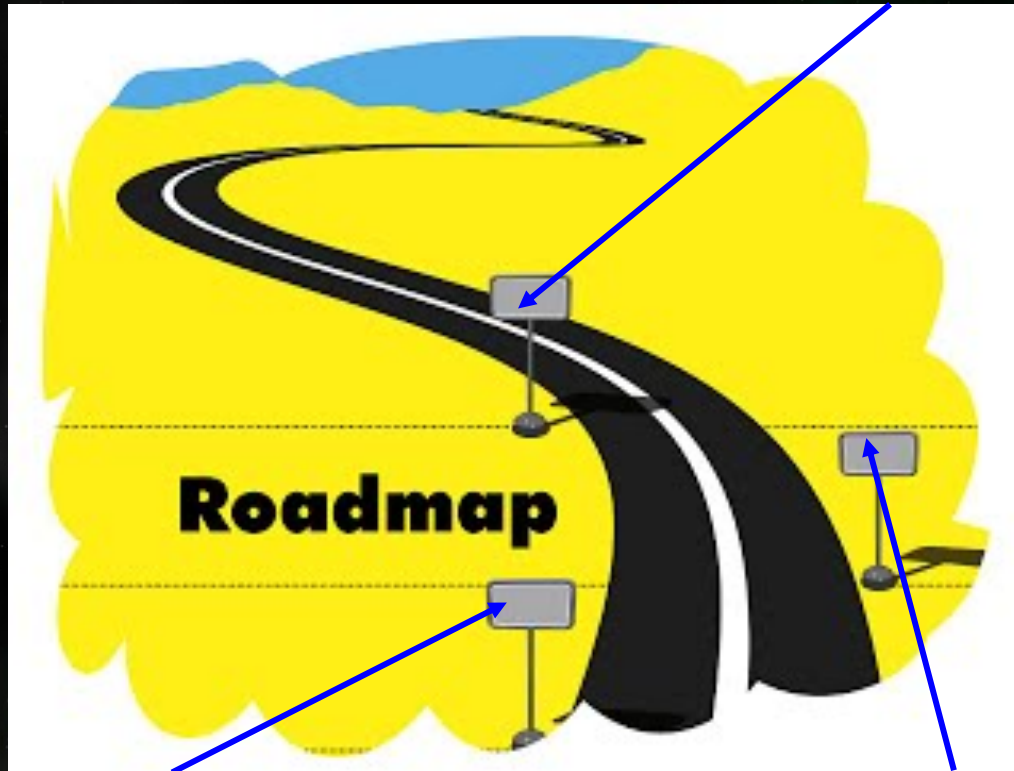
GW170817

Spontaneous/Dynamical
Scalarization

Kent Yagi

Roadmap

Universal Relations



GW170817

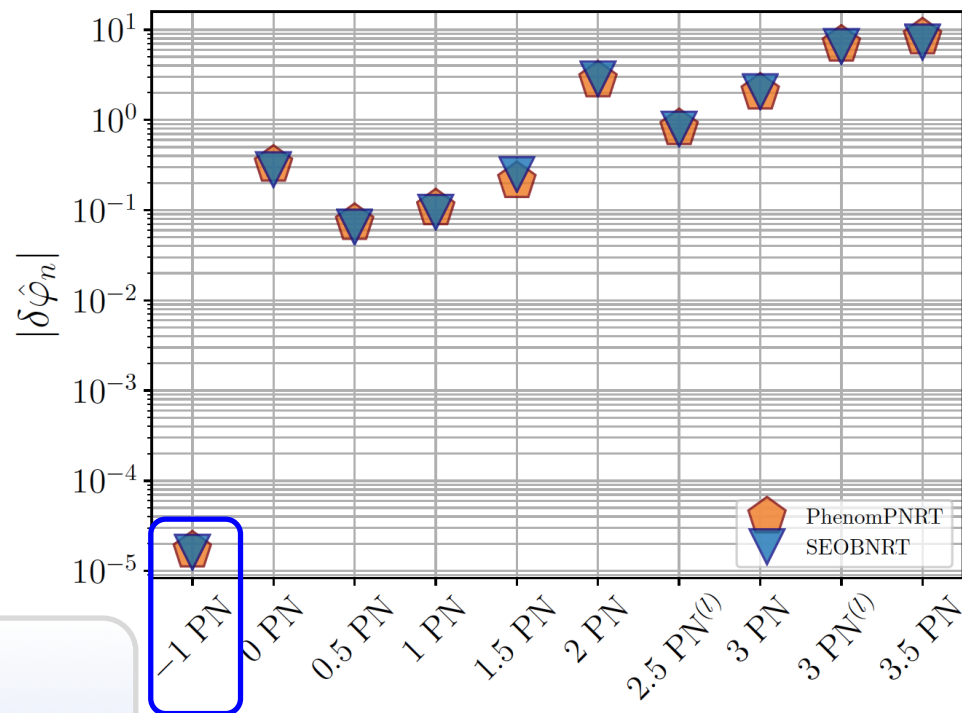
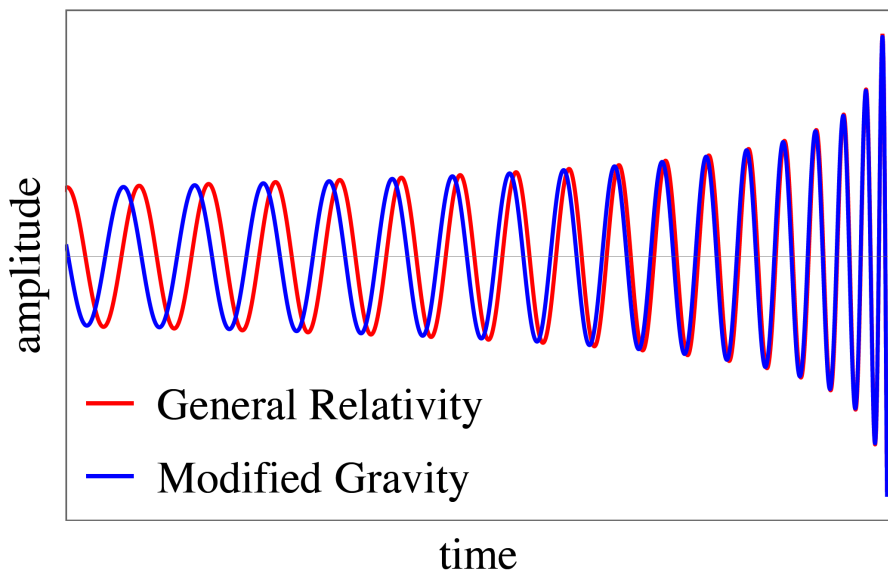
Spontaneous/Dynamical
Scalarization

Kent Yagi

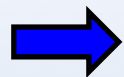
Parameterized Tests

Waveform phase:

$$\Psi(f) = \frac{3}{128\eta} \sum_{k=0}^7 \phi_k^{(\text{GR})} (1 + \delta\phi_k) (\pi m f)^{(k-5)/3}$$



scalar dipole radiation @ -1PN

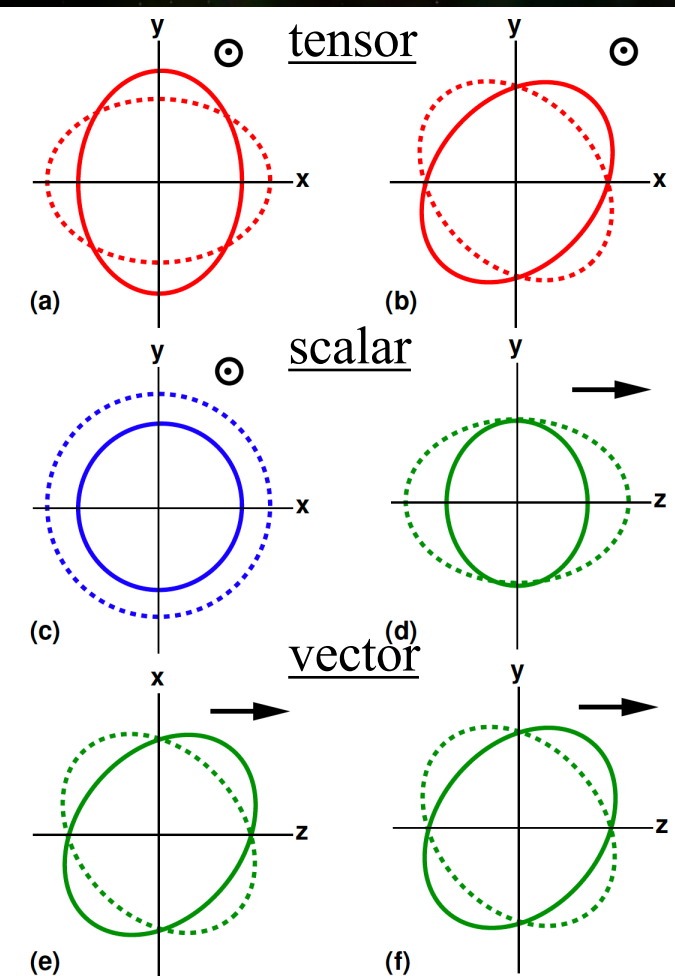


violation of strong equiv. principle

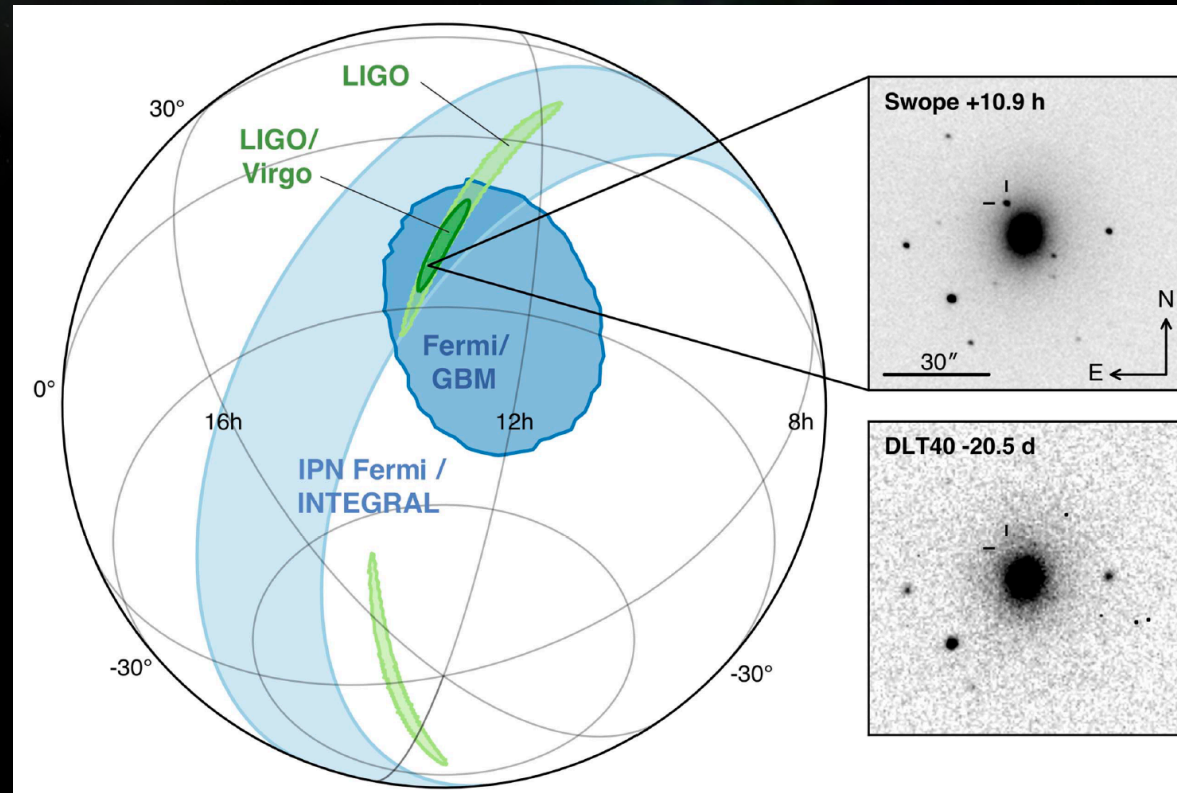
[LVC, arXiv:1811.00364]

Polarization

[LVC, arXiv:1811.00364]



[Will (2014)]



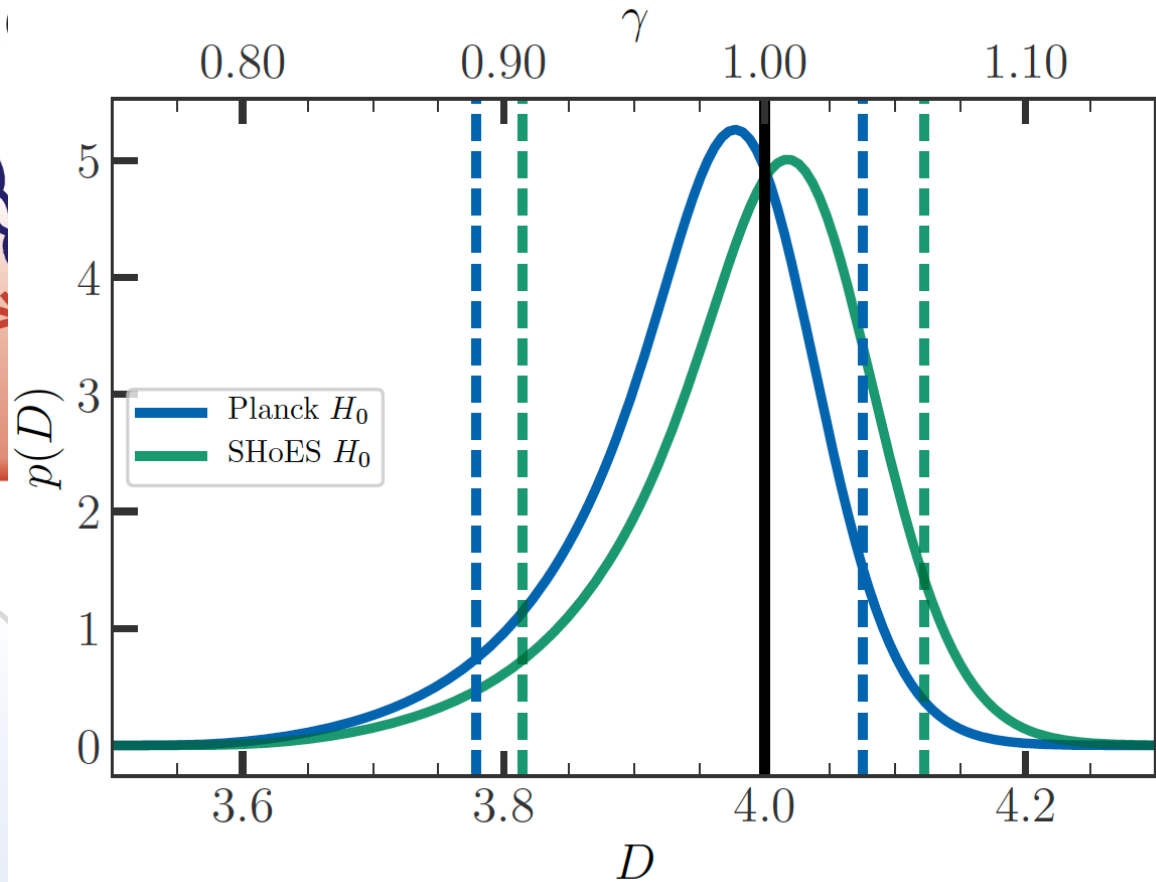
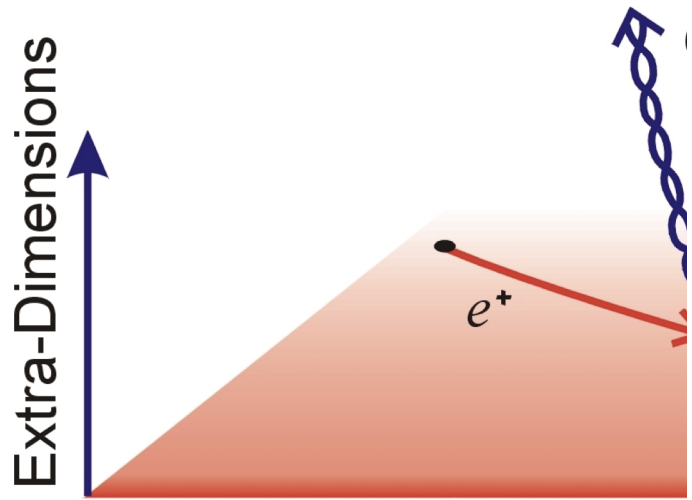
Bayesian Model Selection:

(tensor only) vs (scalar only) = $10^{21} : 1$

(tensor only) vs (vector only) = $10^{23} : 1$

of Spacetime Dimension

[Pardo, Fishbach, Holz & Spergel, arXiv:1801.08160]

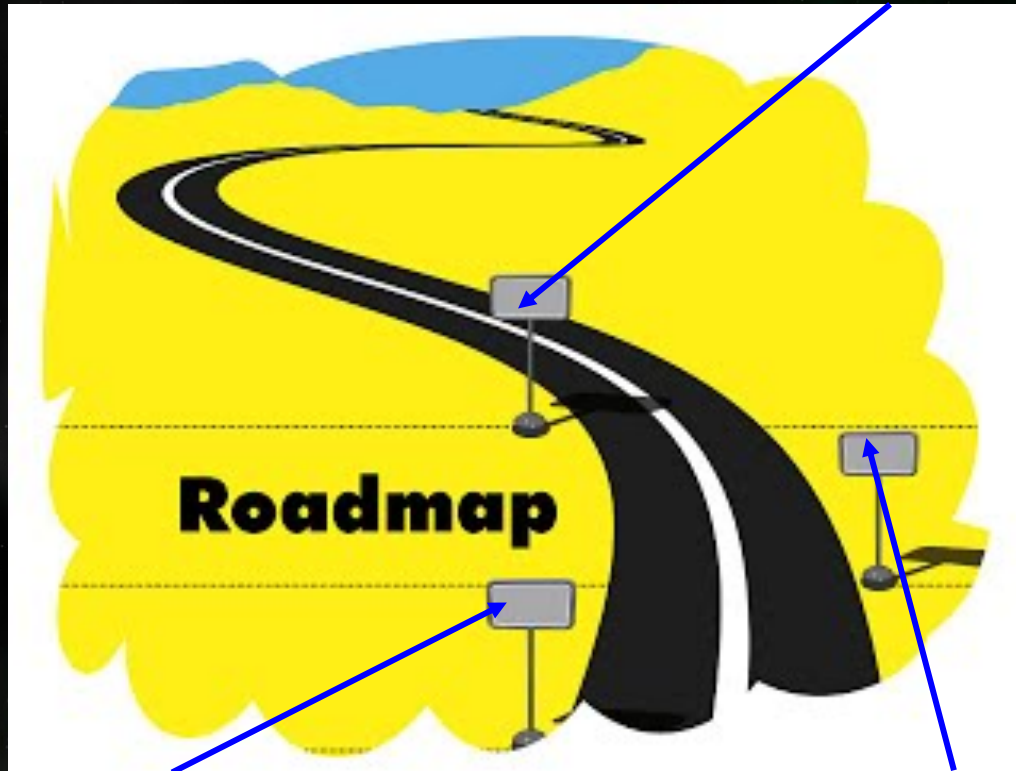


$$h \propto d_L^{-(D-2)/2}$$

 $d_L^{(\text{GW})} \neq d_L^{(\text{EM})}$

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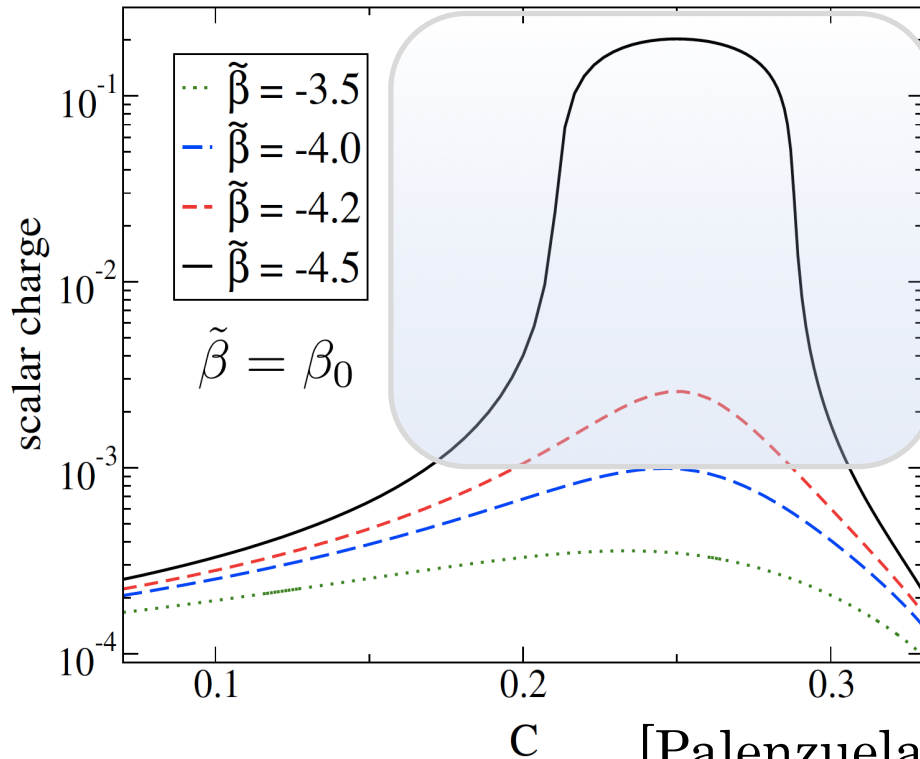
Spontaneous Scalarization

[Damour & Esposito-Farese
(1992, 1993)]

scalar-tensor theories $(g_{\mu\nu}, \phi)$

(α_0, β_0) coupling between the scalar field and matter

controls neutron star scalarization



spontaneous scalarization

(nonlinear growth
of the scalar field)

[Palenzuela et al. arXiv:1310.4481]

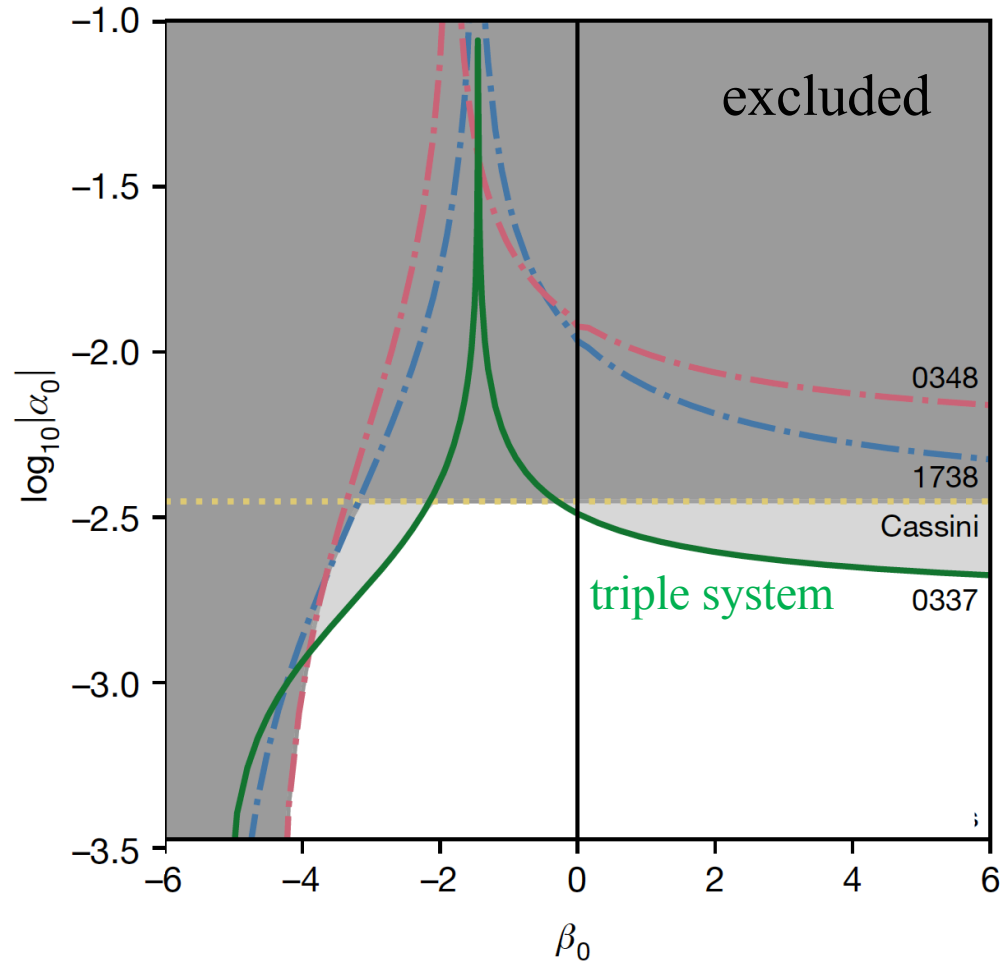
Existing Bounds

[Archibald et al. (2018)]

scalar charge

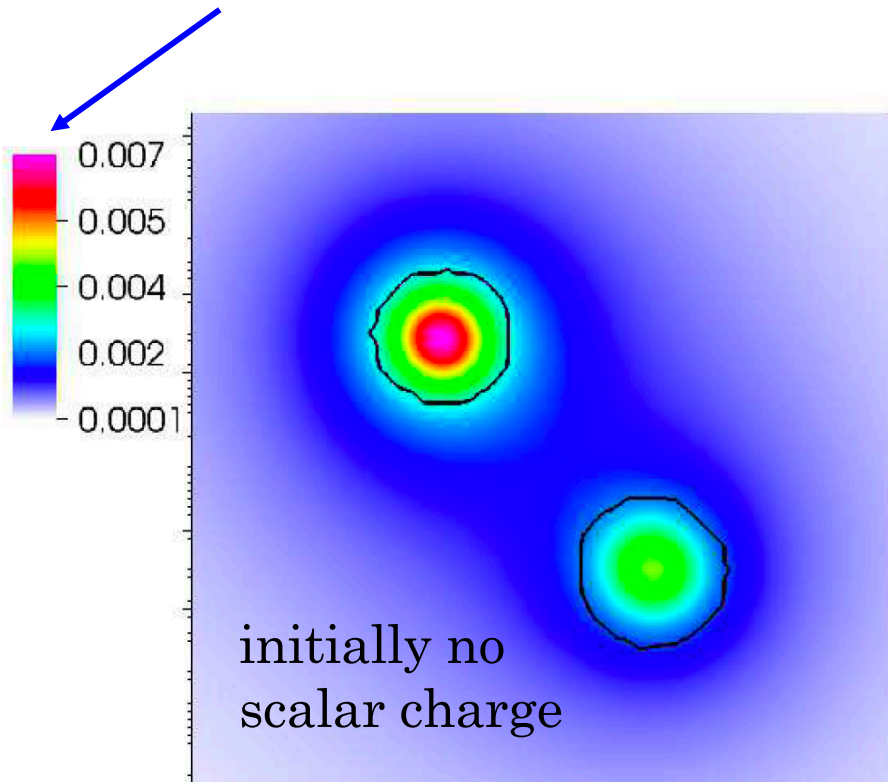


scalar dipole radiation

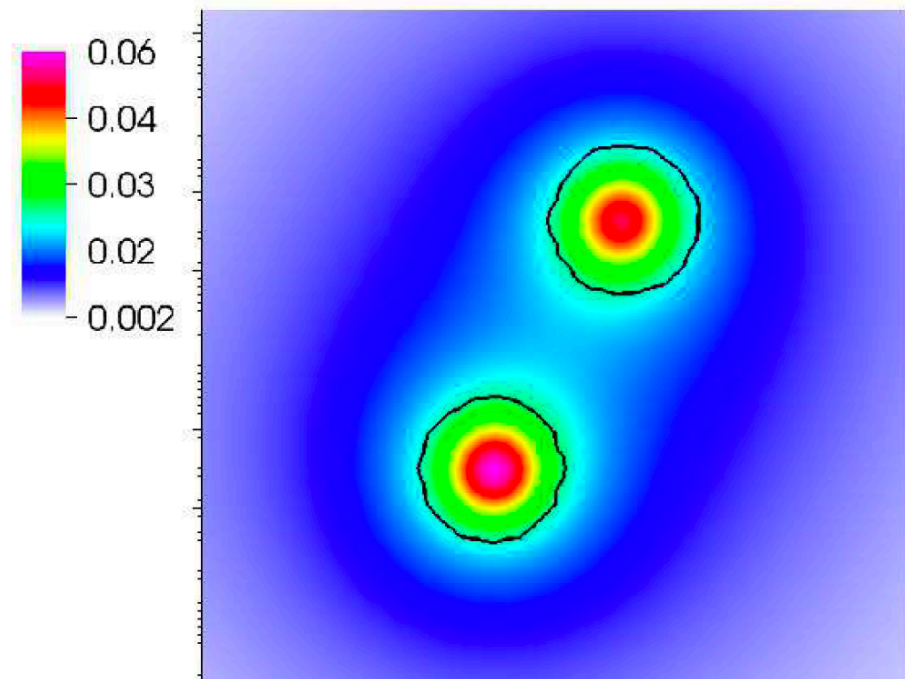


Induced / Dynamical Scalarization

scalar charge

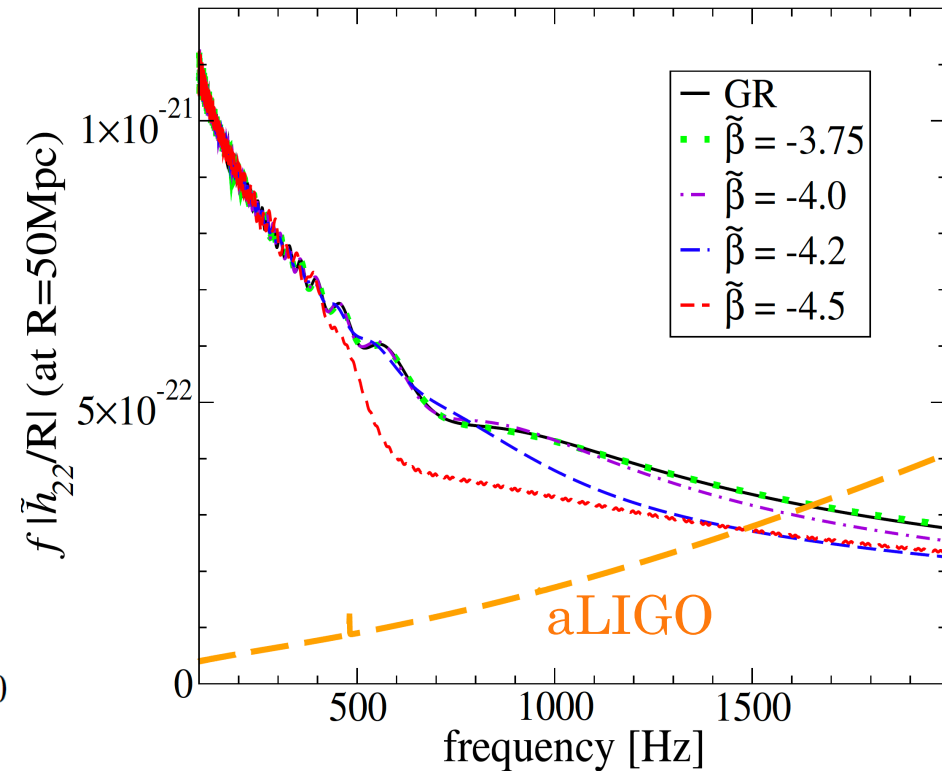
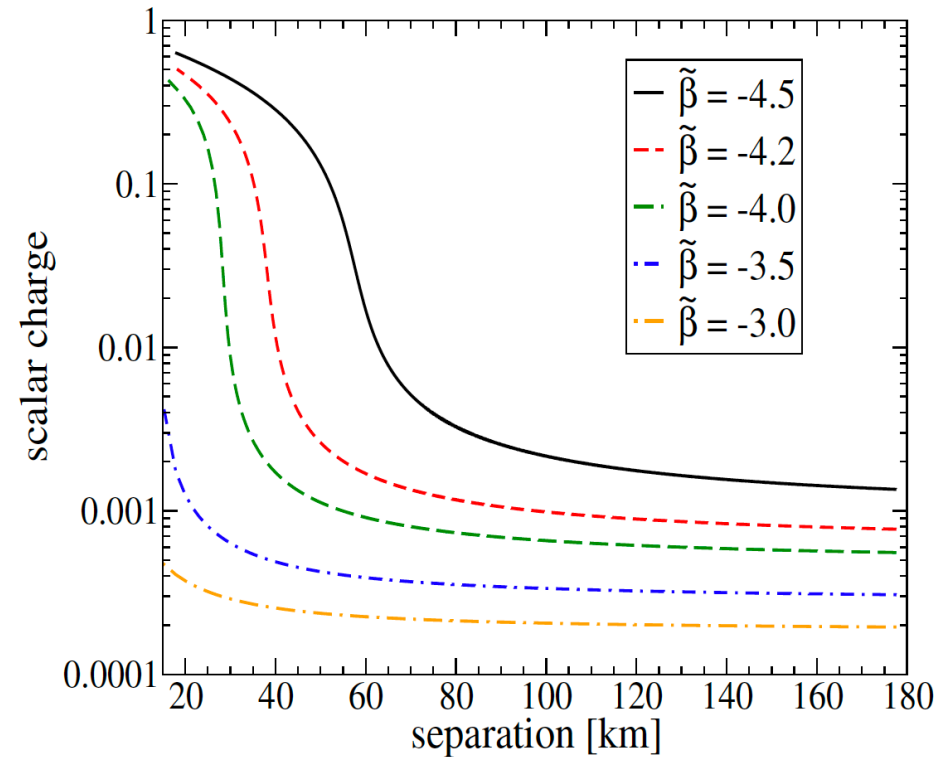


scalar charge induced!



[Barausse et al. arXiv:1212.5053]

Induced / Dynamical Scalarization

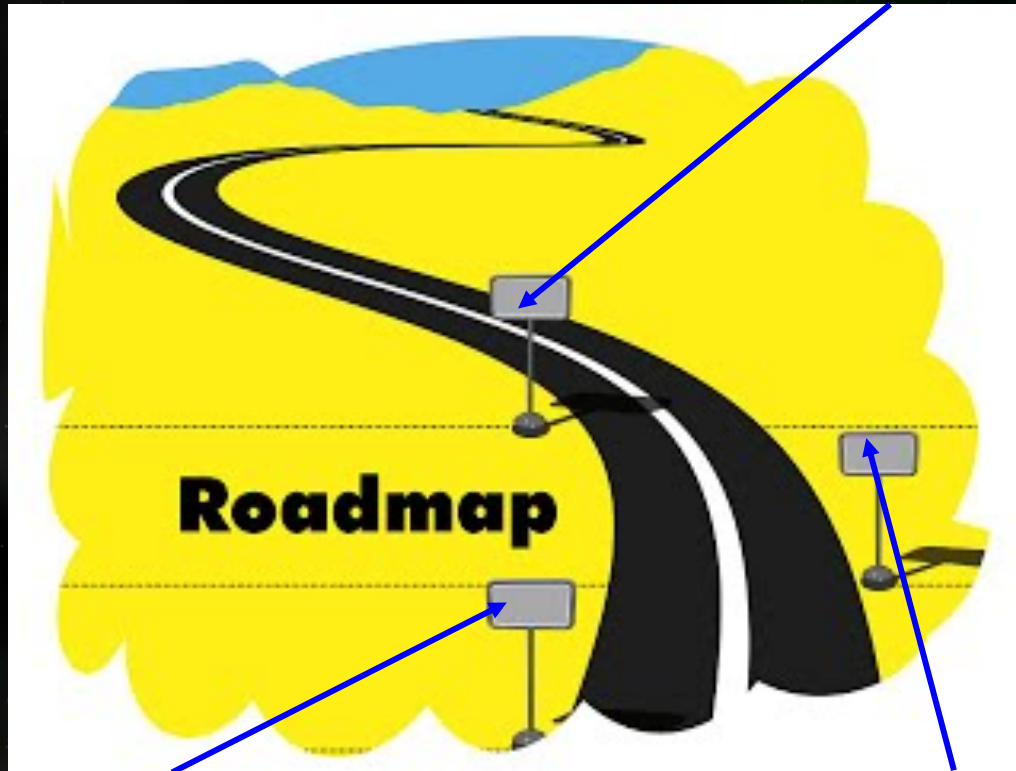


Phenomenon cannot be probed by **binary pulsars!**

[Palenzuela et al. arXiv:1310.4481]

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Tidal Measurement by LVC

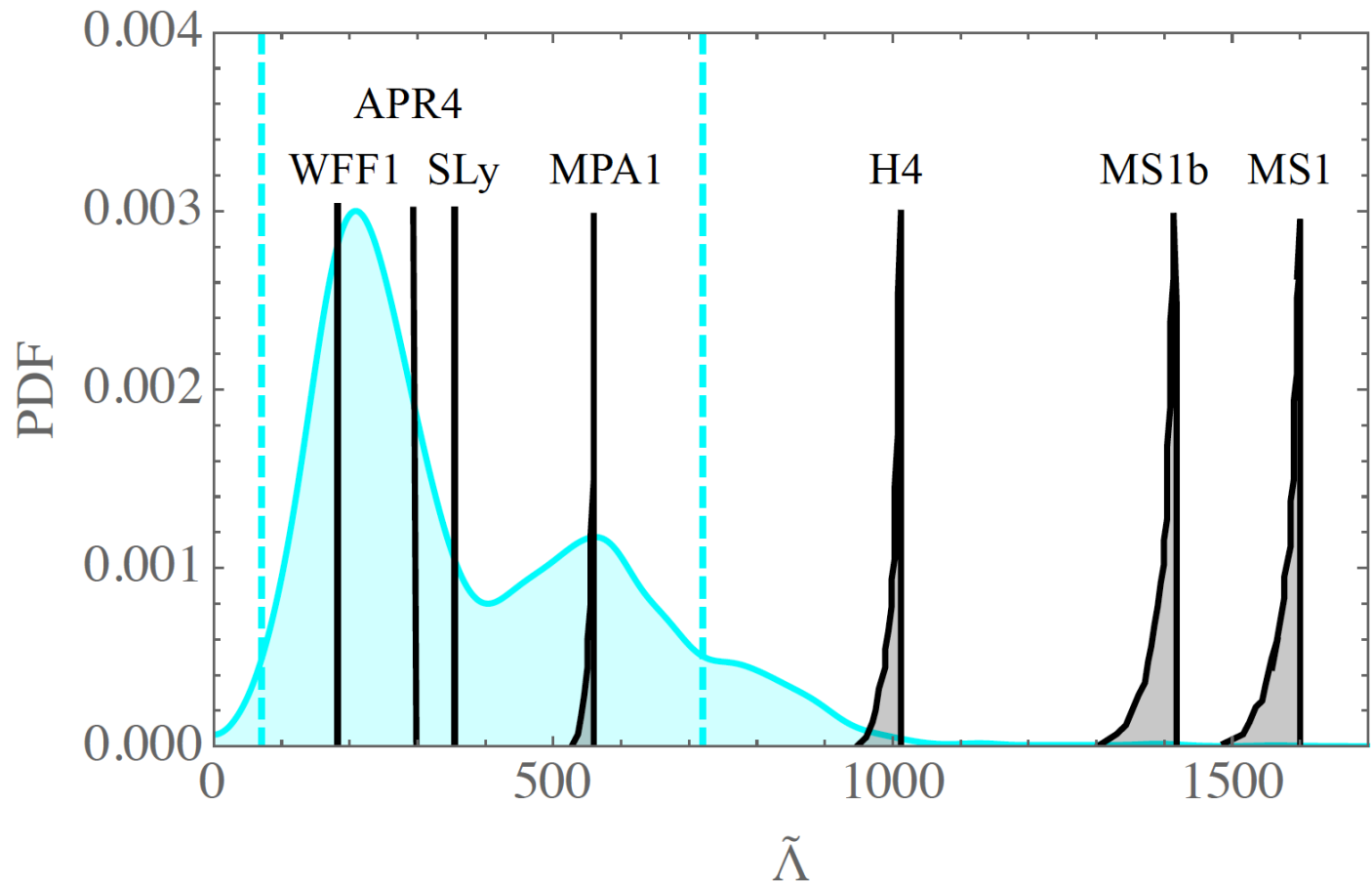
Leading tidal parameter in the waveform:

[LVC, arXiv:1805.11579]

$$\tilde{\Lambda} = \frac{16(1+12q)\Lambda_1 + (12+q)q^4\Lambda_2}{(1+q)^5},$$

tidal measurement can also be used to probe gravity

$$q = m_1/m_2$$

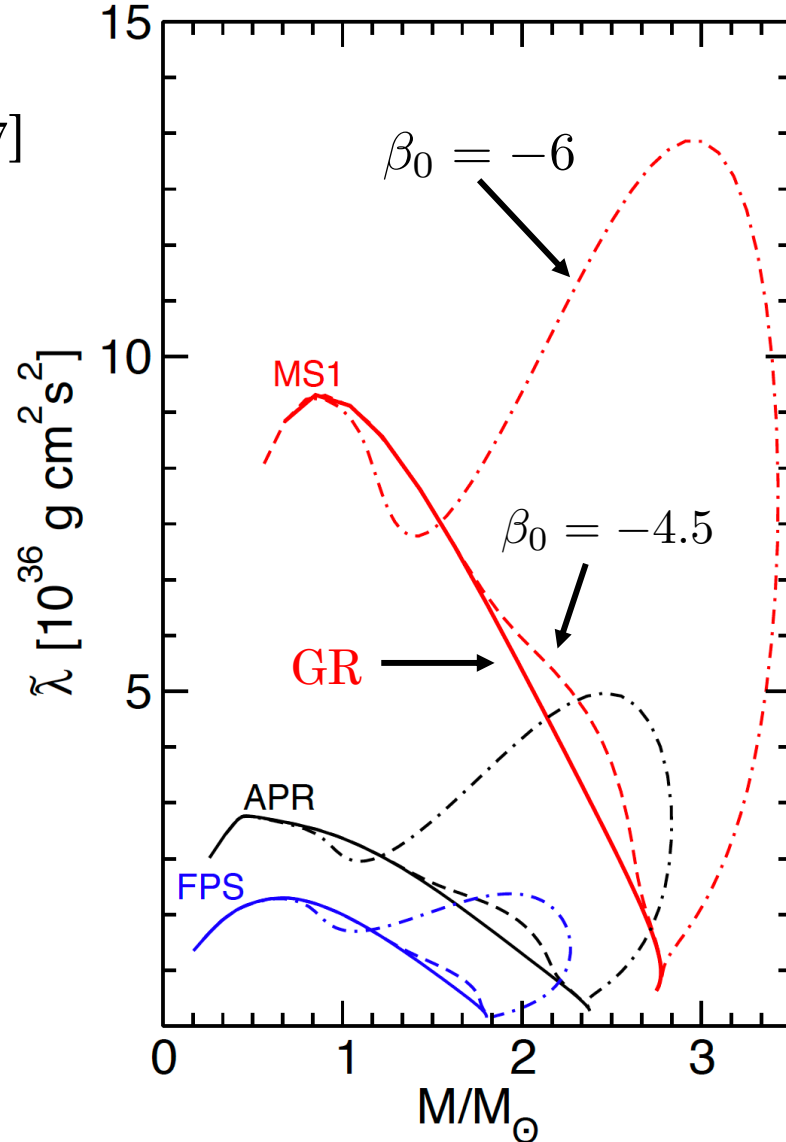


Love numbers in non-GR Theories

scalar-tensor theories

[Pani & Berti arXiv:1405.4547]

EoS vs non-GR...

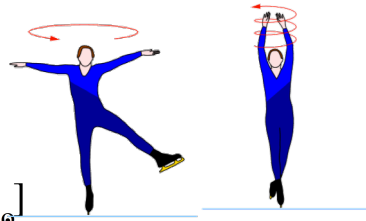


I-Love-Q Relations!

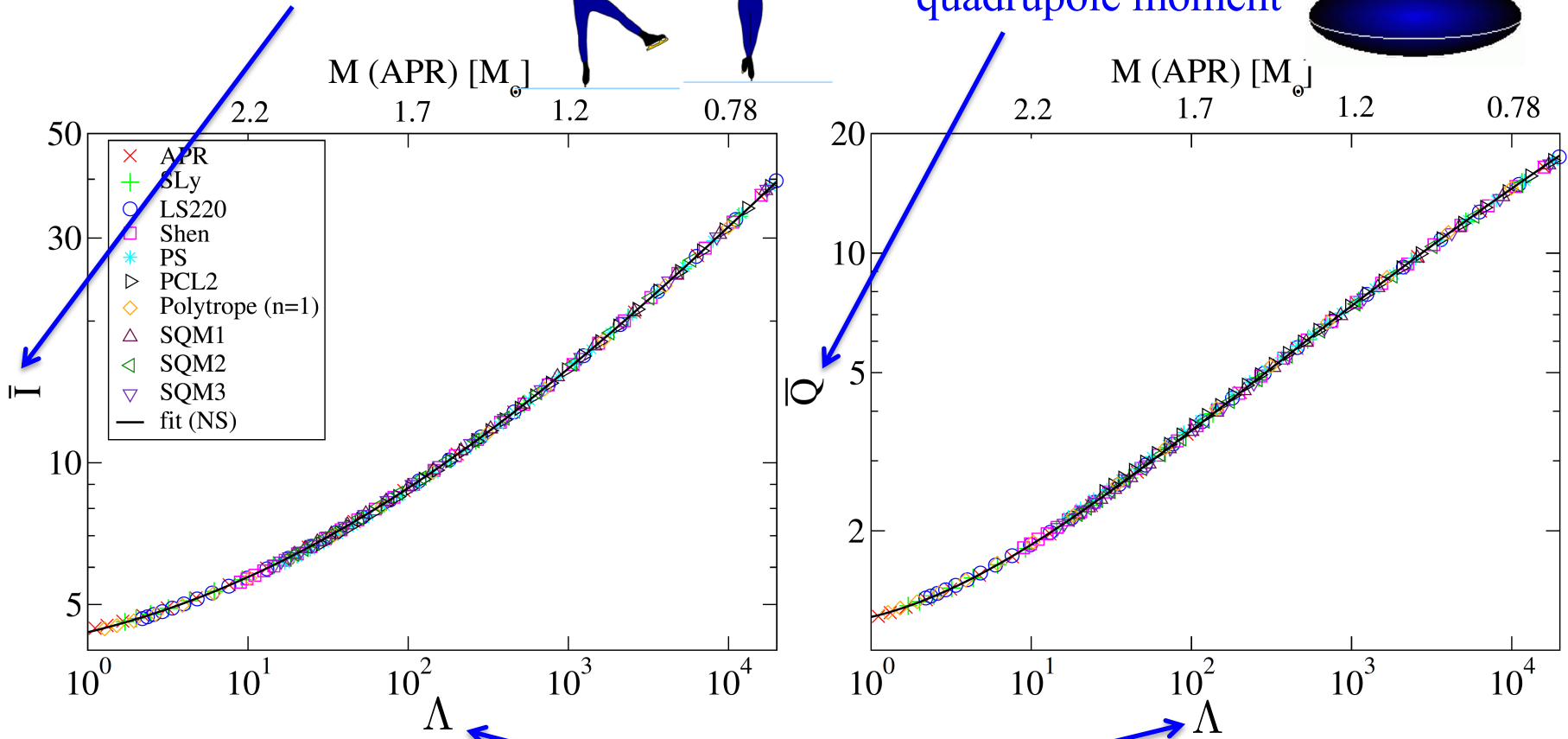
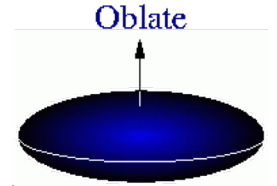
[KY & Yunes arXiv:1302.4499]

[KY & Yunes arXiv:1303.1528]

moment of inertia



(spin-induced)
quadrupole moment



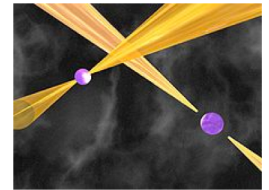
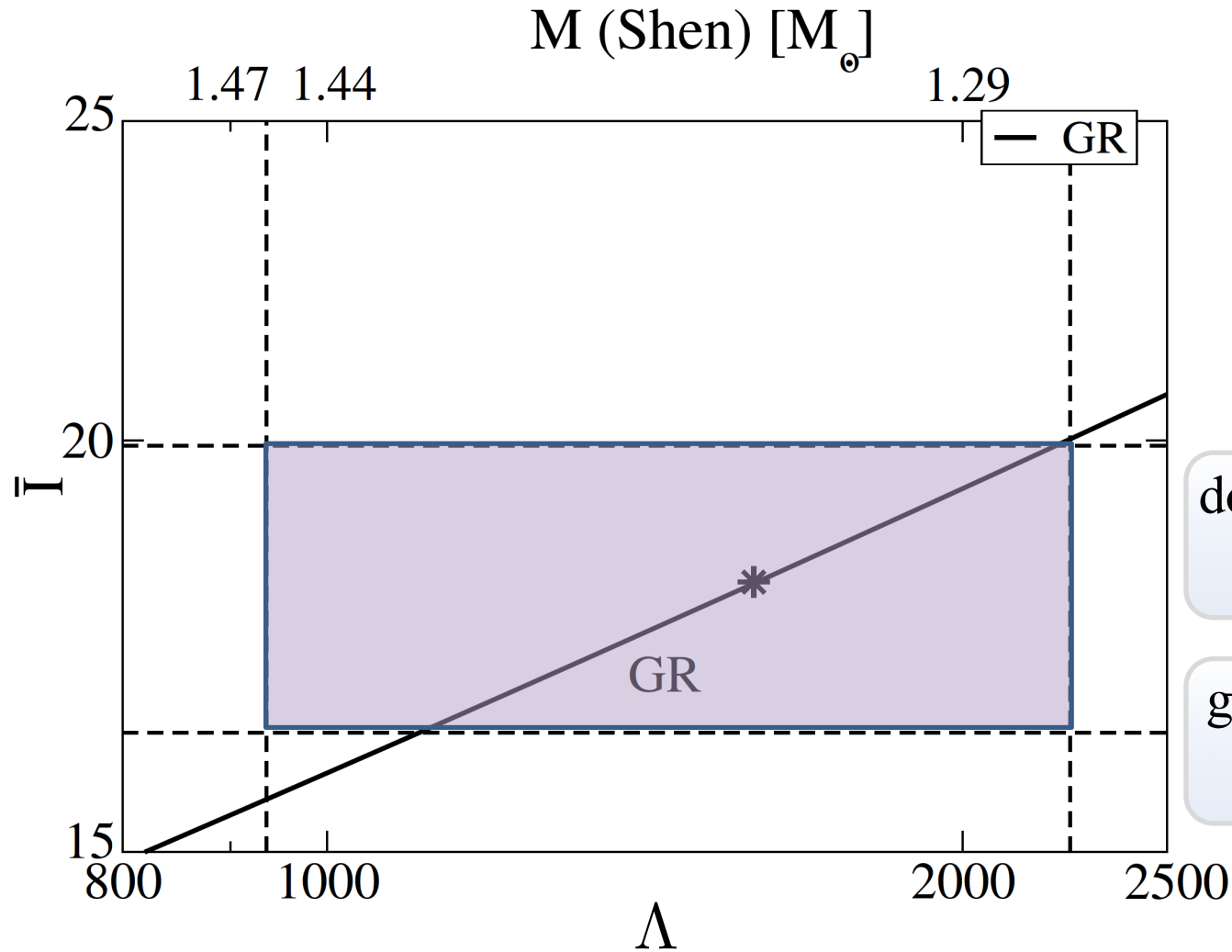
tidal Love number
(tidal deformability)



Strong Gravity Tests

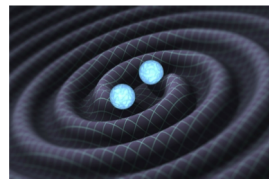
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double pulsar binary
 $\Delta\bar{I}/\bar{I} = 10\%$

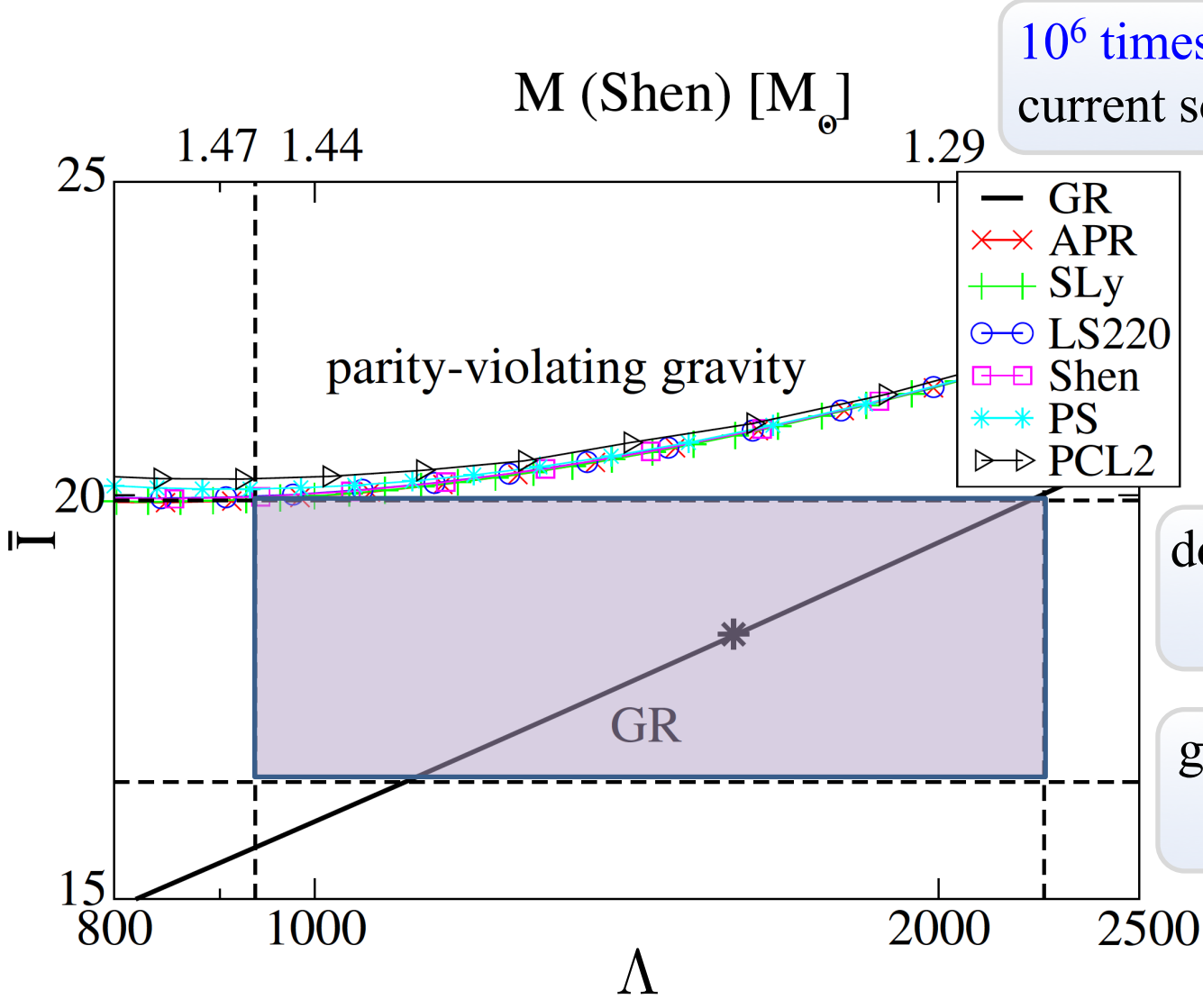
gravitational waves
 $\Delta\Lambda/\Lambda = 40\%$



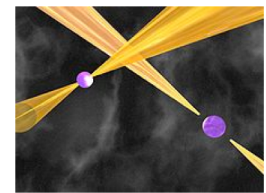
Strong Gravity Tests

[KY & Yunes arXiv:1302.4499]

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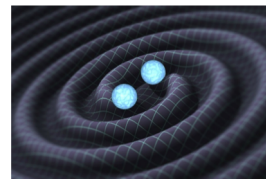


10^6 times stronger than the current solar system bound!



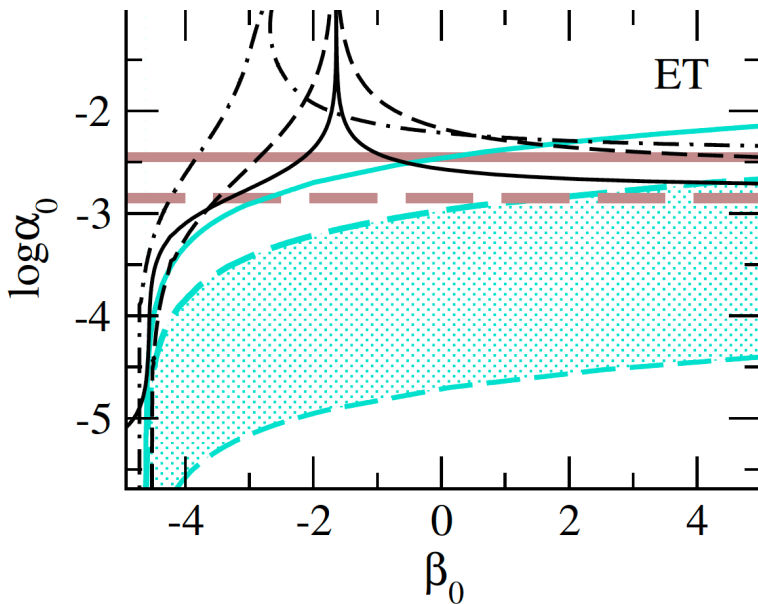
double pulsar binary
 $\Delta \bar{I} / \bar{I} = 10\%$

gravitational waves
 $\Delta \Lambda / \Lambda = 40\%$



Conclusions

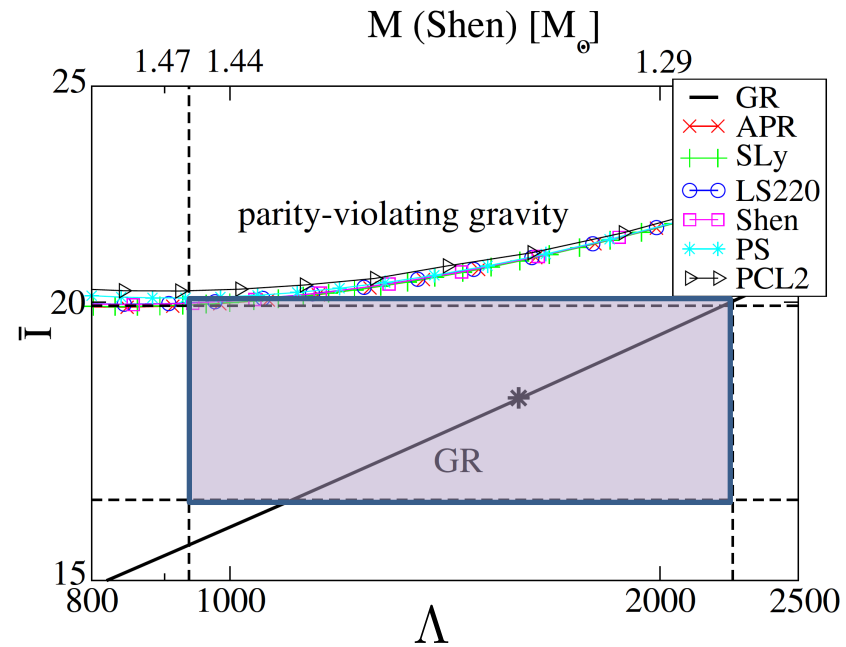
Takeaway



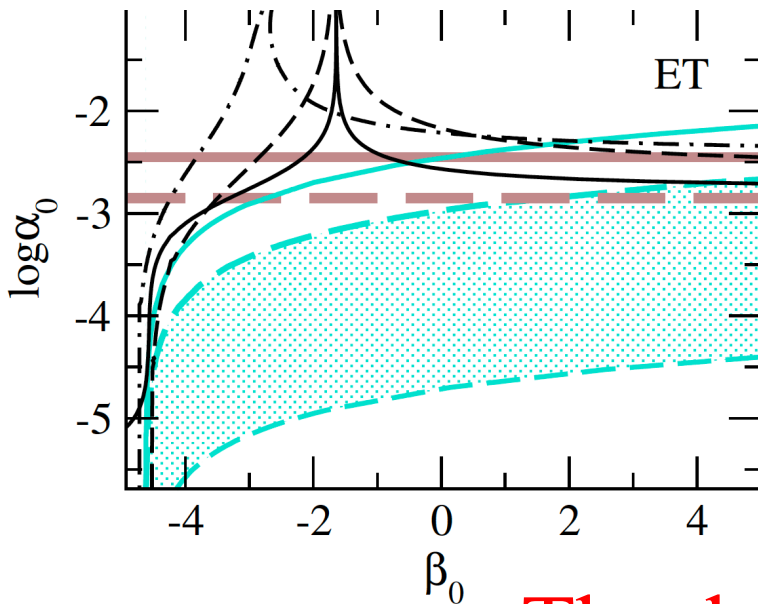
- ✓ stringent bounds on **scalar-tensor** theories expected with **future detectors**
- ✓ **universal relations** can break degeneracies between uncertainties in nuclear physics and grav. physics

✓ GW170817 places stringent bounds on

1. scalar dipole radiation
2. non-GR polarizations
3. # of spacetime dimension



Takeaway



✓ GW170817 places stringent bounds on

1. scalar dipole radiation
2. non-GR polarizations
3. # of spacetime dimension

Thank You!

- ✓ stringent bounds on scalar-tensor theories expected with future detectors
- ✓ universal relations can break degeneracies between uncertainties in nuclear physics and grav. physics

