



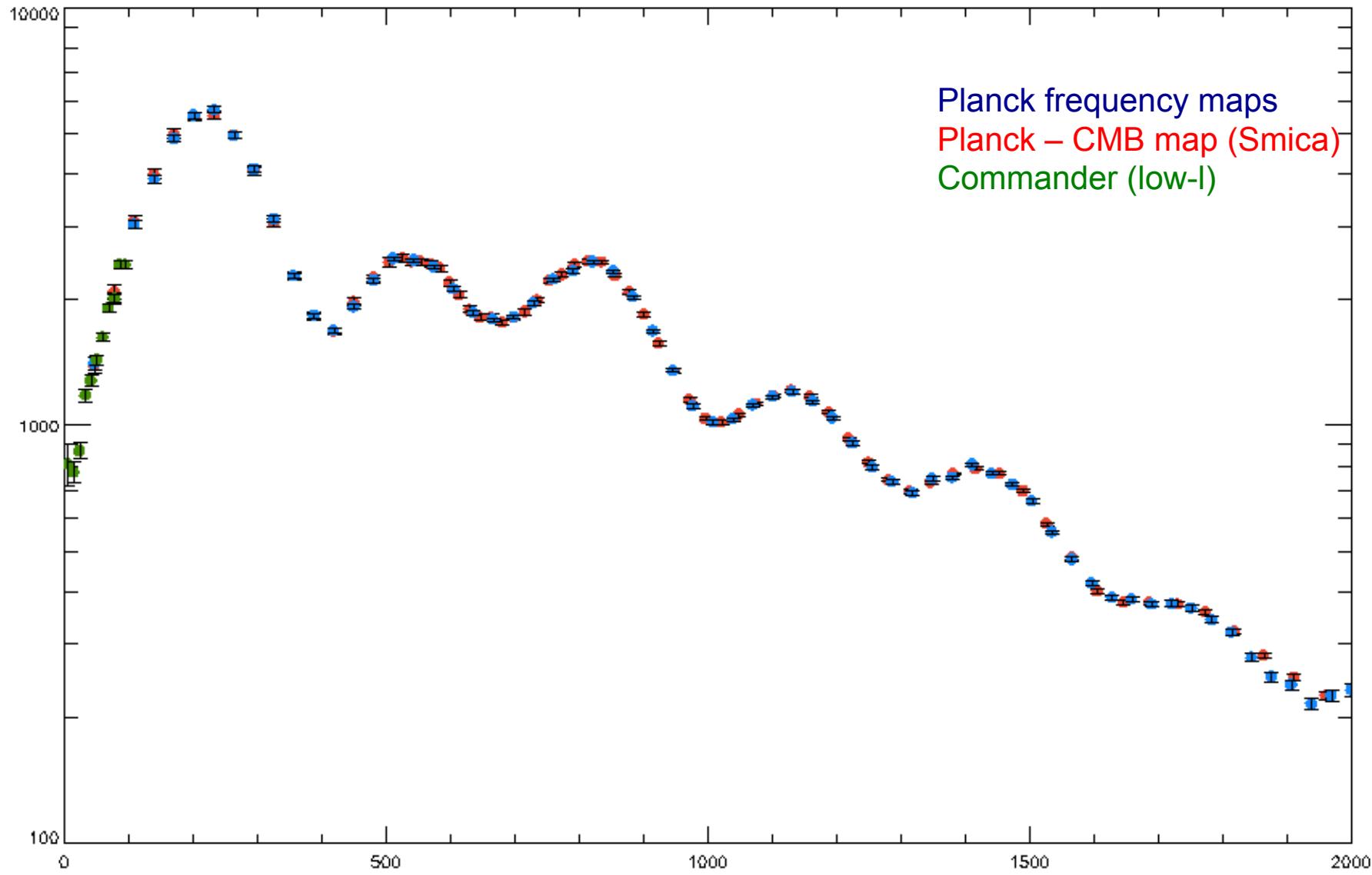
# Implications for Cosmology Comparison with other experiments

Graça Rocha  
JPL/Caltech

On behalf of the Planck collaboration

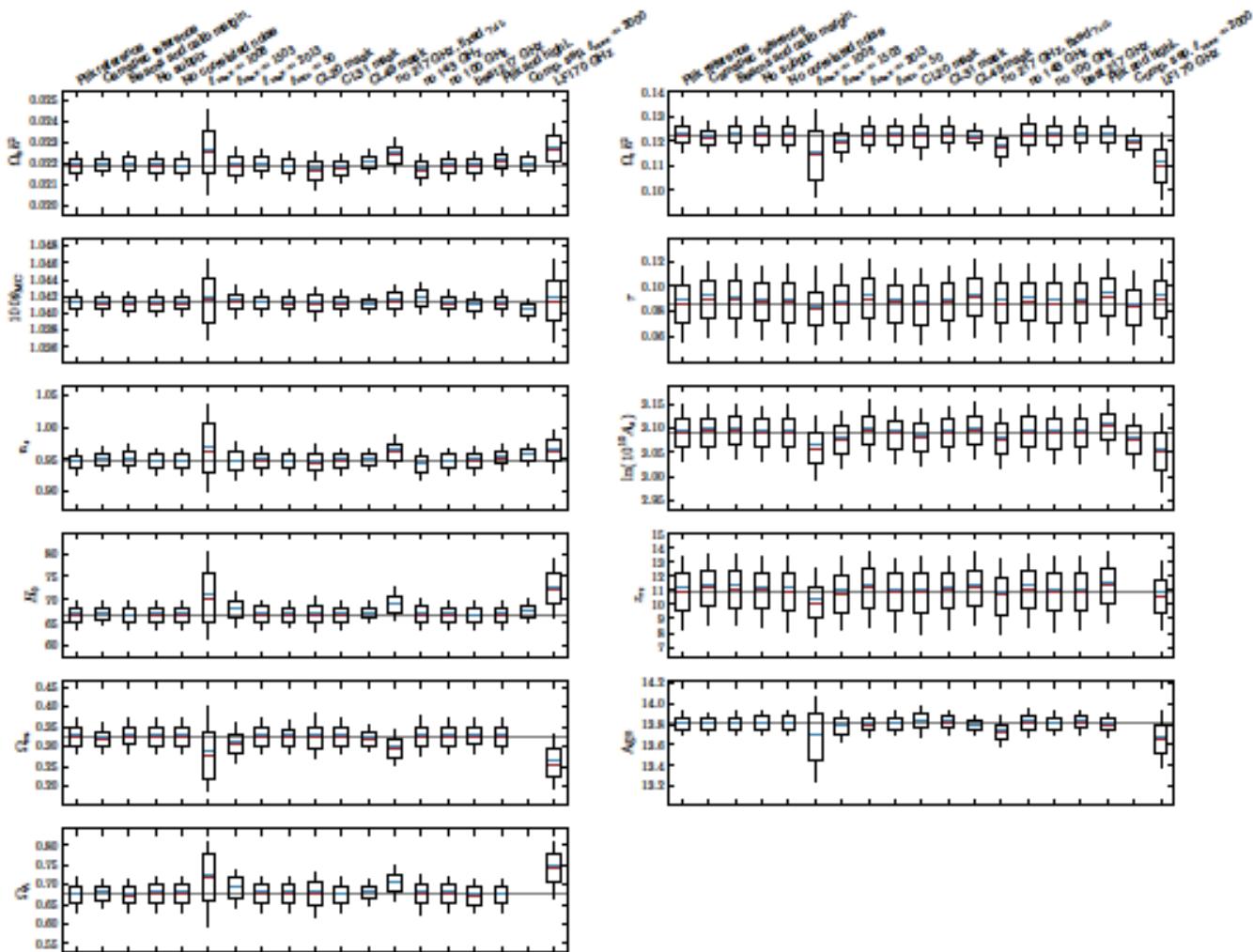


# Planck vs Planck



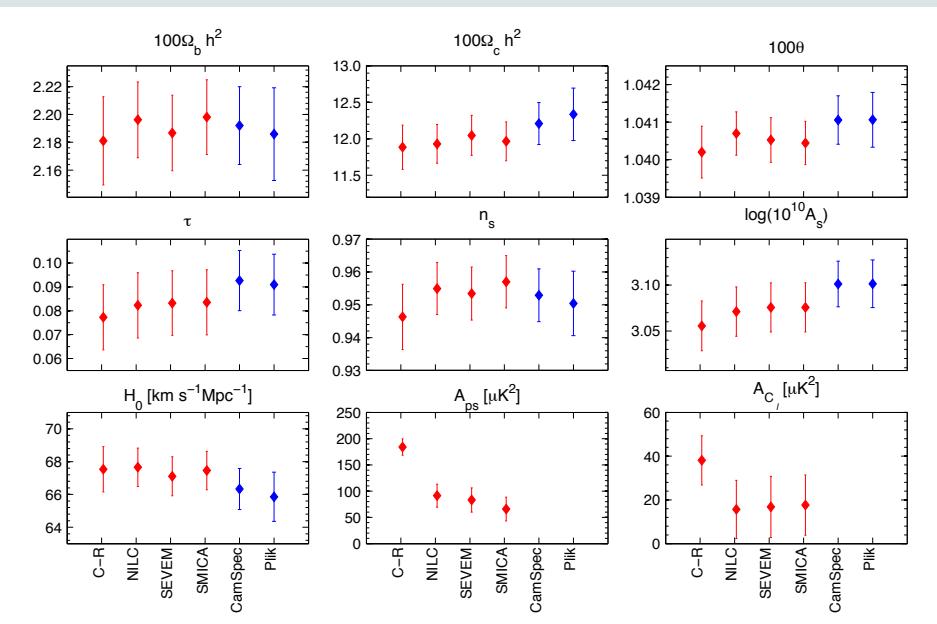
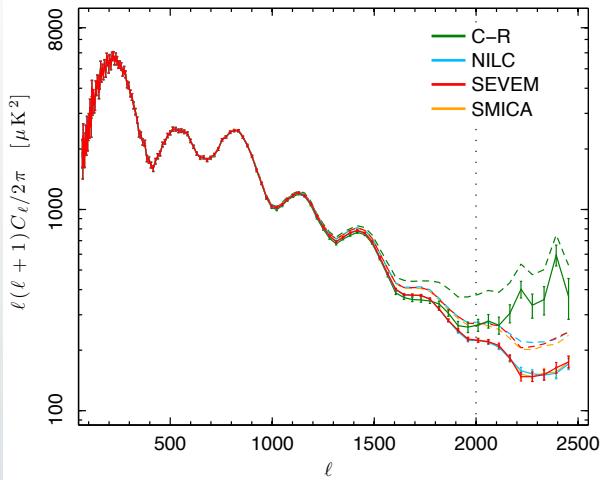
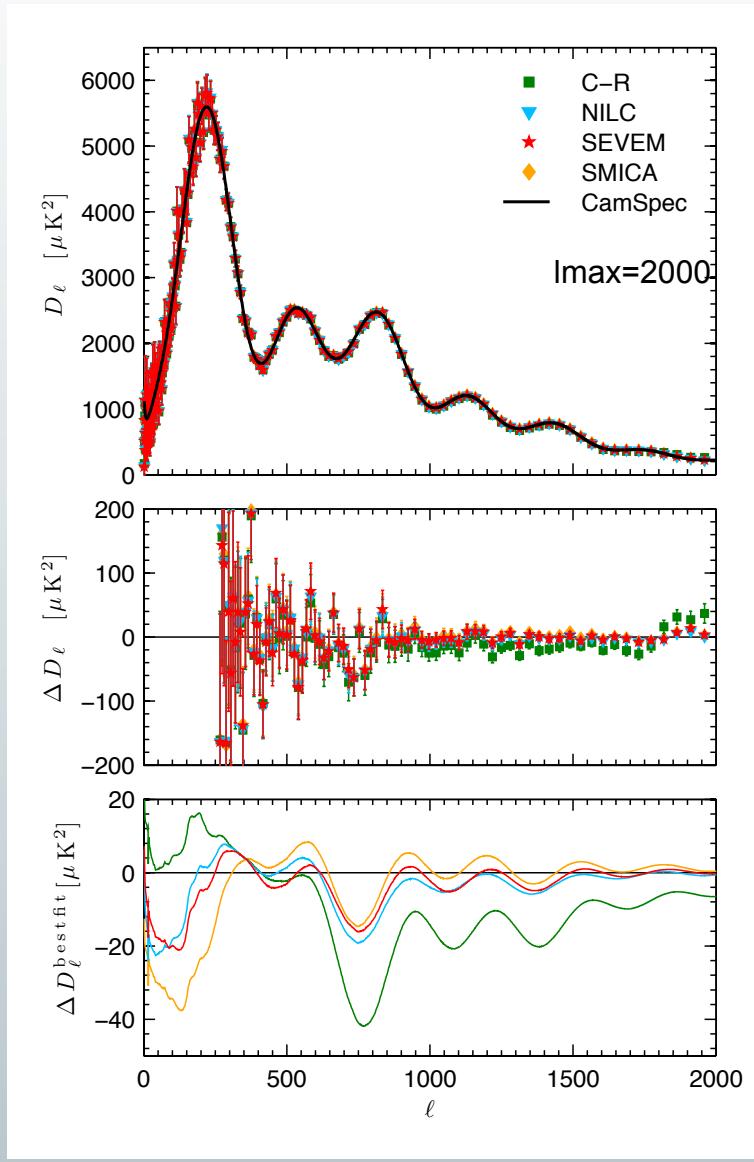
# Planck vs Planck

Planck collaboration: CMB power spectra & likelihood



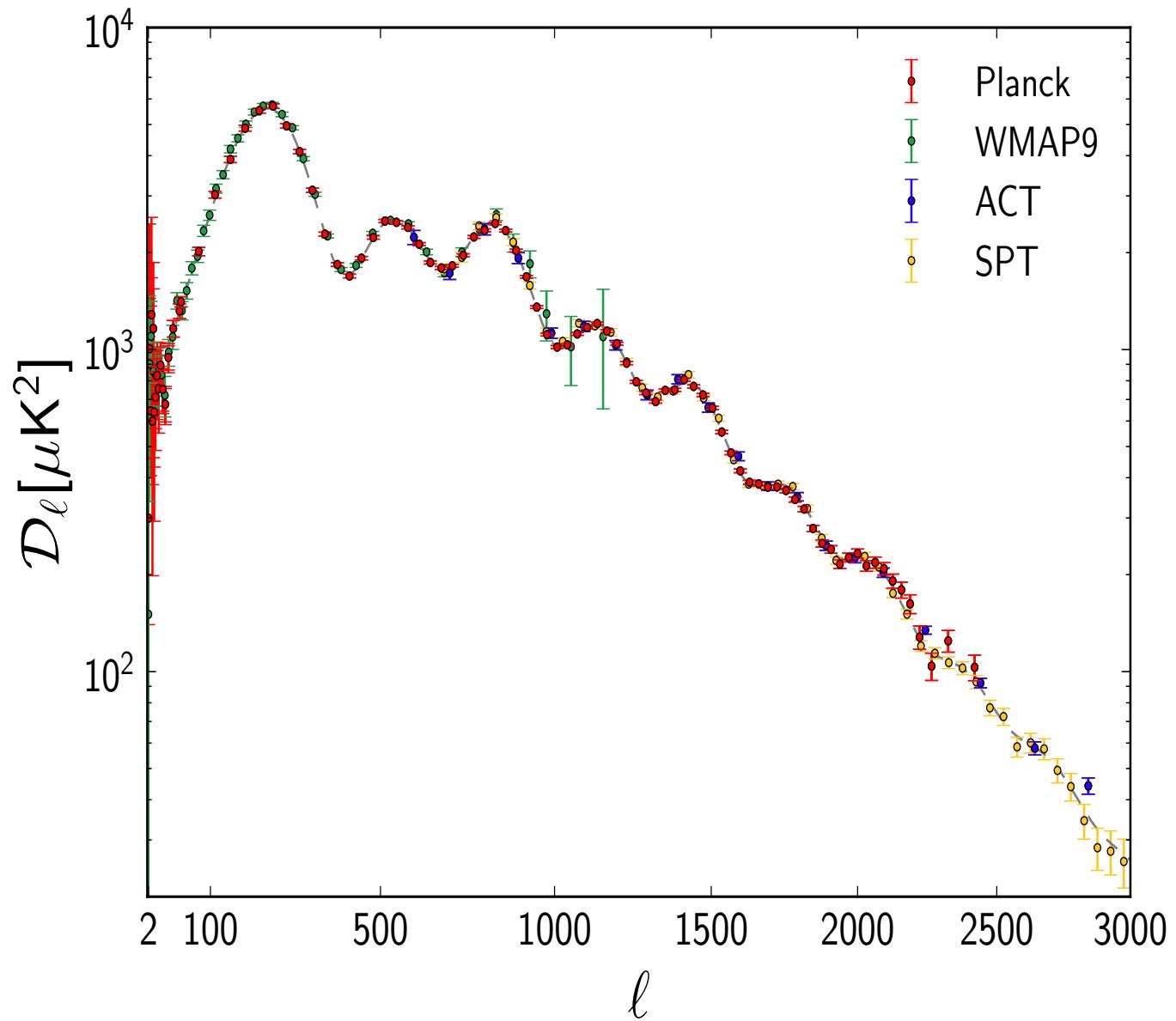
**Figure 24.** Comparison of the distributions of cosmological parameters in the reference case (left) with a set of validation test cases. The red line indicates the median and blue the mean, computed from the posterior histograms. The box shows the 68% confidence interval; the outer line the 95% interval.

# Planck vs Planck





# Planck, WMAP9, SPT, ACT





# Planck vs WMAP9



## $\Lambda$ CDM Parameters, WMAP & Planck

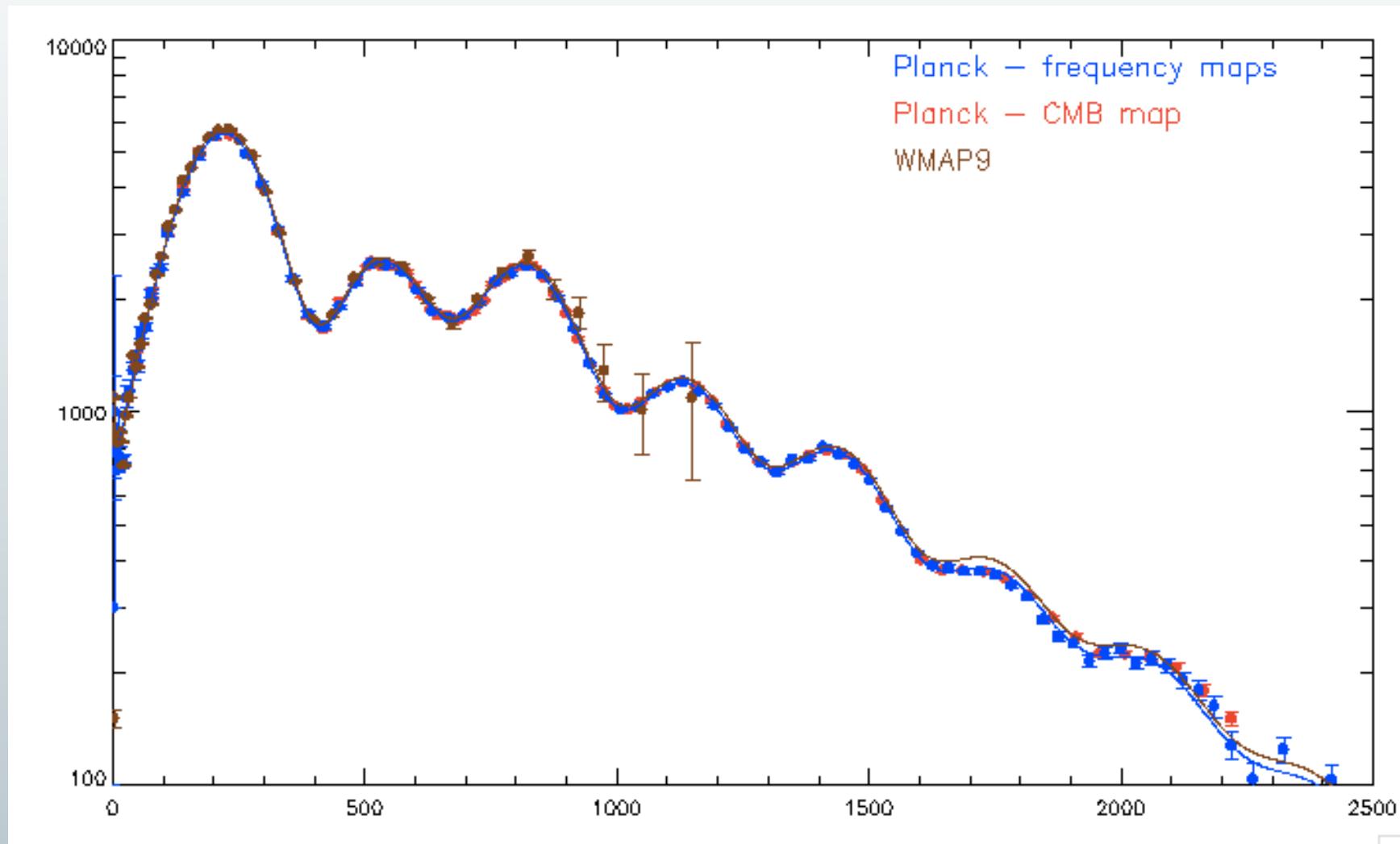
Parameter	Planck ("CMB+Lens")	WMAP (9-year)	Uncertainty ratio (Planck/WMAP)	Shift (WMAP sigma)	Shift (Planck sigma)
Fit parameters					
$\Omega_b h^2$	$0.02217 \pm 0.00033$	$0.02264 \pm 0.00050$	0.66	-0.9	-1.36
$\Omega_c h^2$	$0.1186 \pm 0.0031$	$0.1138 \pm 0.0045$	0.69	+1.1	+1.6
$\Omega_\Lambda$	$0.693 \pm 0.019$	$0.721 \pm 0.025$	0.76	-1.1	-1.4
$n_s$	$0.9635 \pm 0.0094$	$0.972 \pm 0.013$	0.72 <sup>a</sup>	-0.7	-0.97
$\tau$	$0.089 \pm 0.032$	$0.089 \pm 0.014$	2.28 <sup>a</sup>	0	0
$10^9 \Delta_R^2$	...	$2.41 \pm 0.10$ (4.1%)			
$\ln(10^{10} A_s)$	$3.085 \pm 0.057$ (1.8%)	...	0.44 <sup>a</sup>		
Derived parameters					
$t_0$ (Gyr)	$13.796 \pm 0.058$	$13.74 \pm 0.11$	0.53	+0.5	+0.94
$H_0$ (km/s/Mpc)	$67.9 \pm 1.5$	$70.0 \pm 2.2$	0.68	-1.0	-1.5
$\sigma_8$	$0.823 \pm 0.018$	$0.821 \pm 0.023$	0.78	+0.1	+0.13
$100 \theta_*$	$1.04141 \pm 0.00067$	$1.0390 \pm 0.0023$	0.29	+1.0	+3.4

<sup>a</sup>These parameter uncertainties benefit most from polarization data to constrain  $\tau$ .

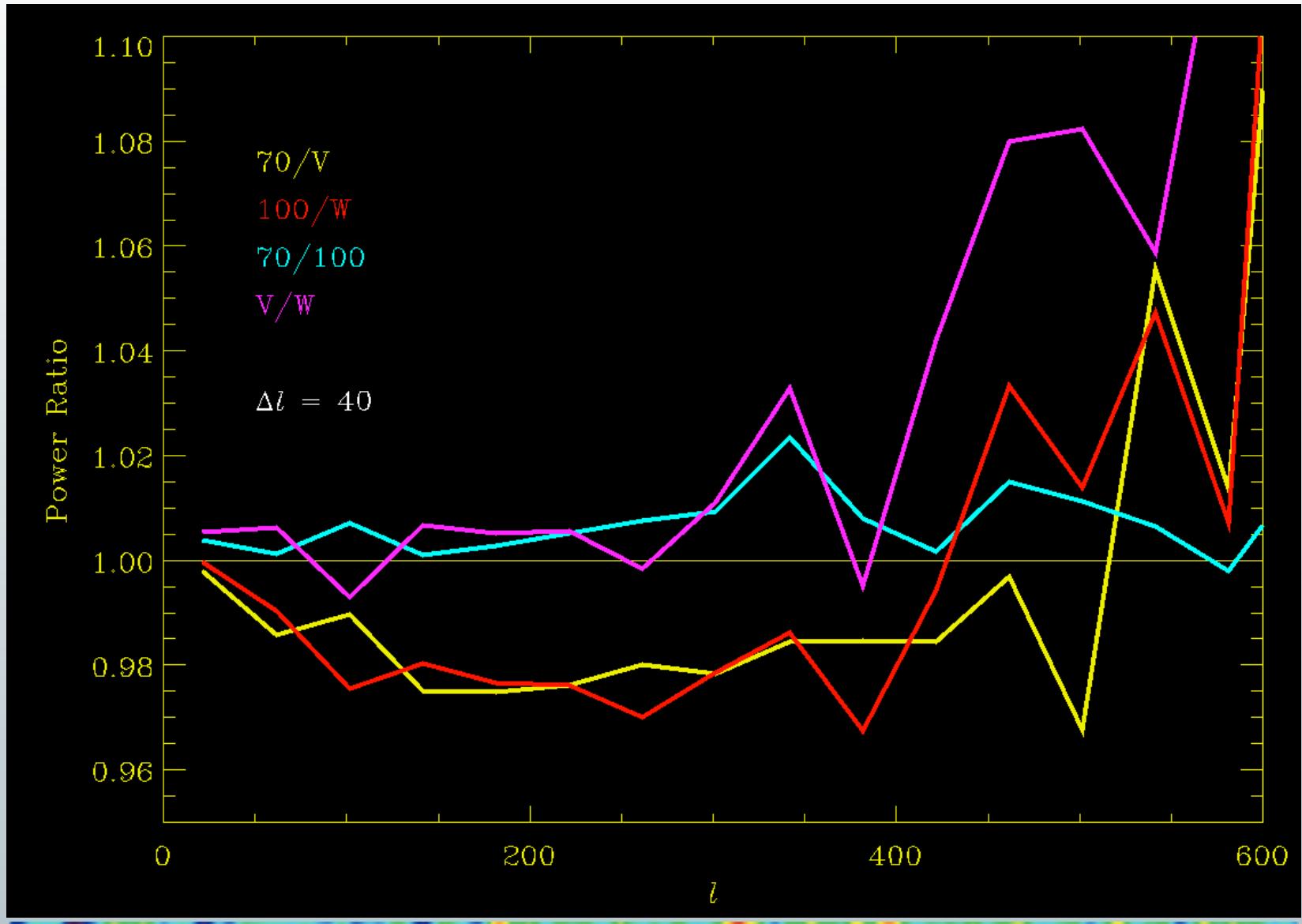
However the shifts in parameters are not independent they talk to the spectra  
the net result is more significative



# Planck vs WMAP9

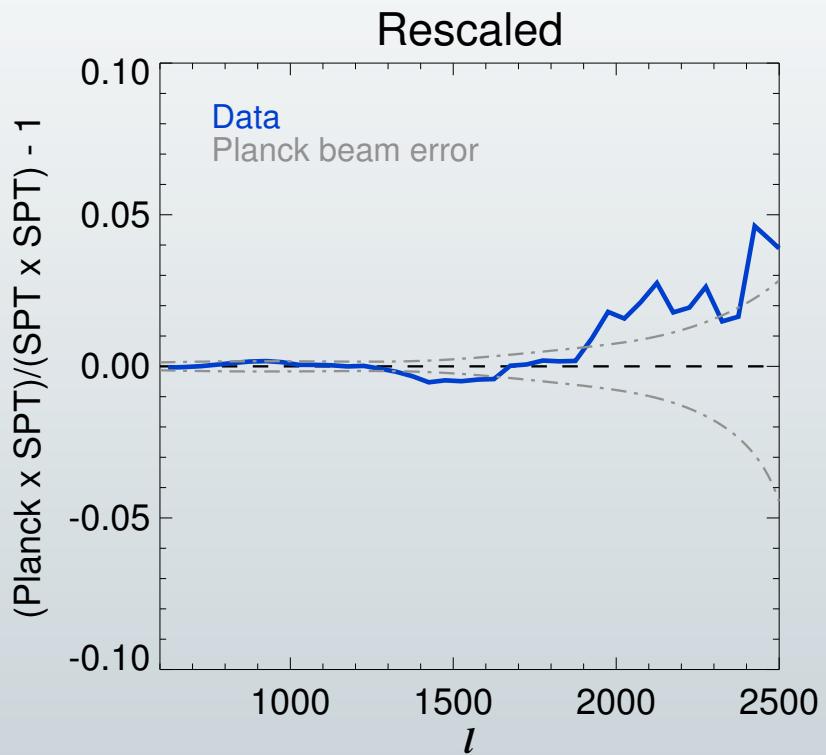
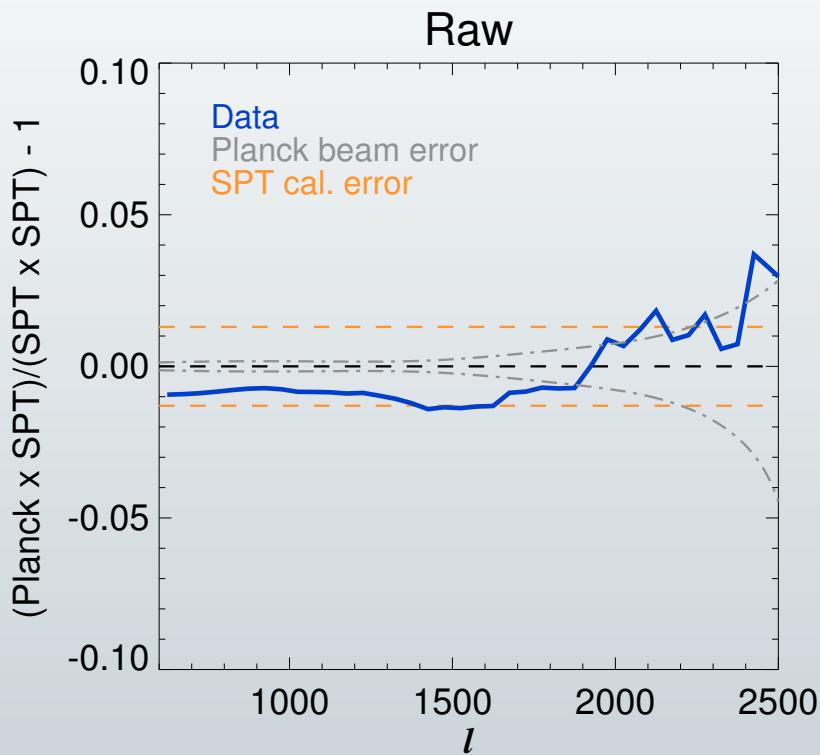


# Planck vs WMAP9





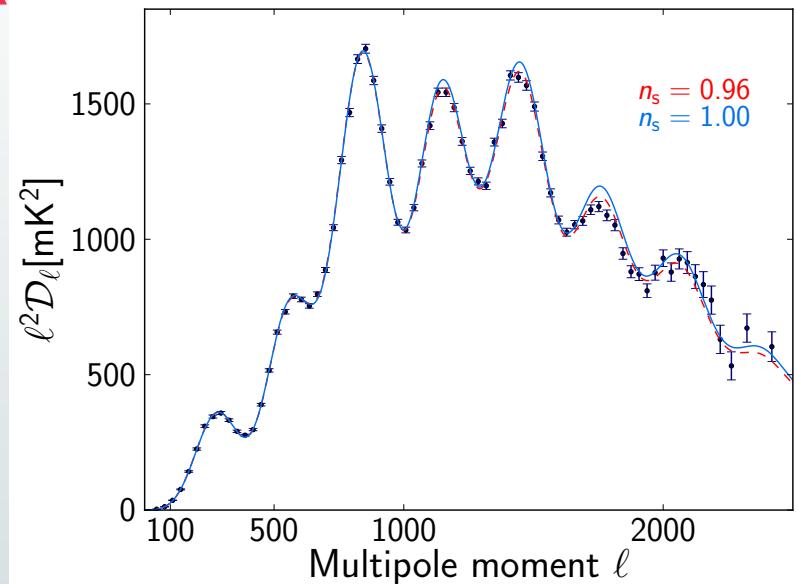
# Planck vs SPT



Courtesy of SPT+Planck team

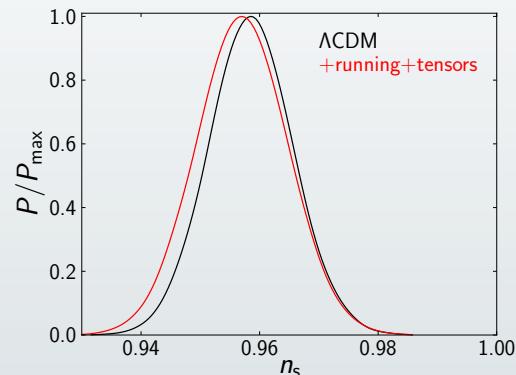
# Extensions to $\Lambda$ CDM model

## Early-Universe physics: $n_s$ , $dn_s/dk$ and $r$



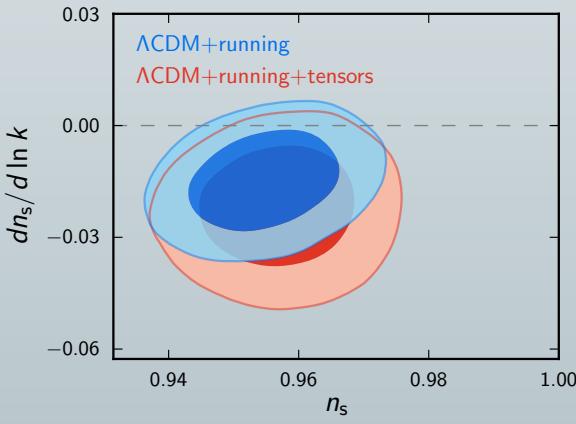
$6\sigma$  departure  
from scale  
invariance

$$n_s = 0.9603 \pm 0.0073$$



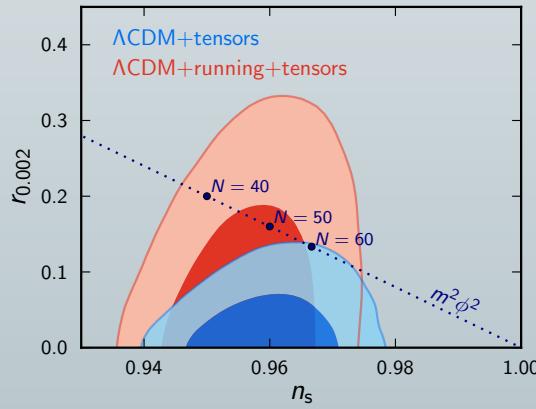
$< 50$

$$dn_s / d \ln k = -0.0134 \pm 0.0090$$



$$r < 0.11$$

$$V = (1.94 \times 10^{16} \text{ GeV})^4 (r_{0.02} / 0.12)$$



$3\sigma$

